MONTHLY WEATHER REVIEW

Editor, EDGAR W. WOOLARD

Vol. 68, No. 12 W. B. No. 1315

DECEMBER 1940

CLOSED FEBRUARY 3, 1941 ISSUED MARCH 24, 1941

THE VARIABILITY OF THE THERMOELECTRIC PYRHELIOMETER FACTOR

By IRVING F. HAND

[Weather Bureau, Washington, D. C., March 1940]

For several years pyrheliometers utilizing copper-constantan thermopiles have been used by the Weather Bureau at Washington, D. C., and the Blue Hill Observatory of Harvard University at Milton, Mass., to measure normal incidence radiation. ¹ In every case the readings of the thermoelectric pyrheliometers were checked against readings of substandards, chiefly the Smithsonian silverdisk and the Marvin resistance pyrheliometers, in order to determine factors by which to multiply scale readings to obtain radiation values in gram-calories. After finding that there is a change in these factors with radiation fluctuations, we began a series of comparisons late in 1938 between our substandards and the thermoelectric pyrheliometer; and a new and longer series was commenced in March 1939, after the recording micromax potentiometer had been thoroughly adjusted by factory experts. Comparisons also were made between our substandard pyrheliometers and a vacuum thermoelectric pyrheliometer² registering on both a micromax potentiometer and an eye-read microammeter.

The appreciable errors, introduced by the change in resistance of the elements in the vacuum thermopile with temperature variations, induced us to change from the measurement of current to the null potentiometric method. This change is appreciable because of the relatively large ratio of the resistance of the couple to that of the total circuit; that is, the resistance of the couple is 7 ohms as compared with 8 ohms of the microammeter and less than 1 ohm of the leads, while 7 ohms is the maximum resistance which we can introduce externally and still retain proper scale deflections.

Only 14 series of comparisons were made between the substandard pyrheliometers and the vacuum thermocouple recording on a microammeter, and these give a probable error of ± 4.5 percent when a single mean for a full calorie range is used as a constant factor. By drawing a line of best fit through the plotted readings, the probable error is reduced to ± 2.7 percent. An attempt was made to determine the effect of free-air temperature changes, but without success.

Unquestionably the effect of the Stefan-Boltzman law enters into the cause of the varying factors; but calculations from available data fail to give results comparable with the line of best fit, and it is thought, therefore, owing to lack of sufficiently precise data on the characteristics of the alloys used, the dimensions, and other quantities, that the only practical method of obtaining the factors is through a long series of direct comparisons.

June 1924.

Single junction vacuum normal incidence pyrheliometer made by Leland B. Clark, of the Astrophysical Observatory of the Smithsonian Institution, Washington, D. C.

A much longer series of 337 comparisons was made between our substandard pyrheliometers and the Eppley normal incidence pyrheliometer, and 298 comparisons between the same substandard instruments and the Clark vacuum pyrheliometer, both recording on a potentiometer. Table 1 lists all comparisons and corresponding factors for both instruments; figure 1 shows a plot of the mean factors, as abscissas, against radiation values in gram calories as ordinates, for the Eppley pyrheliometer.

In the case of the Eppley pyrheliometer the probable error of a single observation from the line of best fit in the range 0.85-1.45 gram calories is ± 0.37 percent, and the probable error of the means of a series of 10 is ± 0.24 percent. However, if the mean value for all observations is used for a constant factor, the probable error of individual readings from this constant factor is ± 1.18 percent for the same range, but somewhat larger for the entire range ordinarily covered when making normal incidence measurements from air-mass 5.0 to as close to 1.0 air mass as is practicable.

We would expect the probable error of a series to be less than that of a single observation, because radiation receipt never is uniform. Moreover, the thermoelectric records are continuous, whereas the substandard pyrheliometers give readings only every minute or every 4 minutes, depending upon the type used.

The probable errors of both instruments with various combinations are tabulated in table 2.

Table 1.—Determination of factors by which to multiply scale readings on Leeds and Northrup potentiometer to obtain normal incidence radiation in gram calories

Date and hour		Scale	Factor	Scale	Factor	Means			
angle	Q	(Ep- pley)	Q/Ep- pley	(Clark)	Q/ Clark	Q	Eppley	Clark	
1939	,								
March 3: 3:24	1. 134 1. 133 1. 118 1. 123 1. 134	46. 0 46. 0 45. 0 45. 5 46. 0	. 0247 247 246 246 247			1. 128	0. 02466		
	1. 130 1. 115 1. 133 1. 151 1. 161	46. 0 45. 0 46. 0 46. 5 46. 5	248 248 246 248 248			1. 138	. 02472		
3:10	1, 167 1, 171 1, 197 1, 206 1, 207	47. 0 47. 0 48 . 0 49. 0 48. 5	248 249 249 246 249			1. 190	. 02482		
1:21	1. 405 1. 389 1. 383 1, 380	55, 5 54, 5 54, 0 54, 0	253 255 256 256						
1:16	1. 389	55.0	253			1.389	.02546		

¹ The first instance of this method of pyrheliometric measurement known to the writer was described by Ladislaus Gorczyński in the Monthly Weather Review, 52:299-301, June 1924.

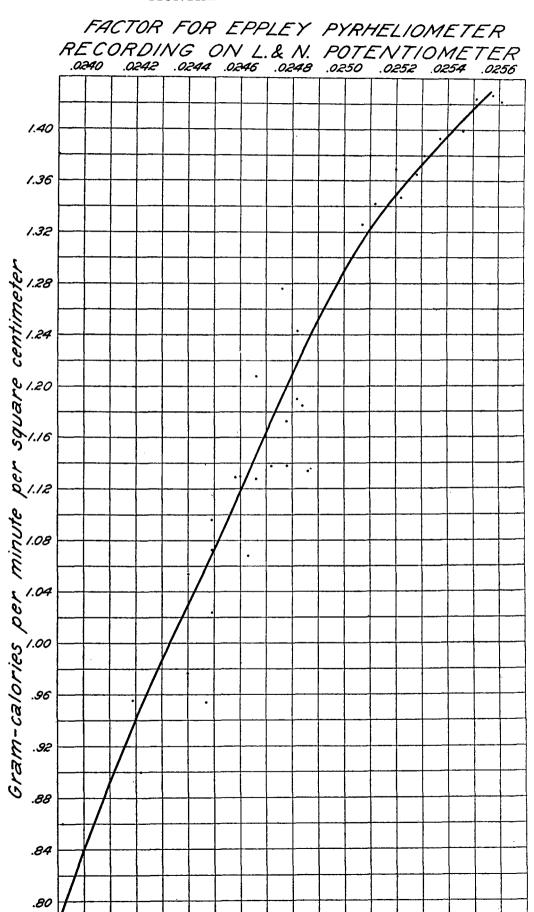


FIGURE 1.—Plot of the mean factors, as abscissas, against radiation values in gram-calories as ordinates, for the Eppley pytheliometer.

Table 1.—Determination of factors by which to multiply scale readings on Leeds and Northrup potentiometer to obtain normal incidence radiation in gram calories—Continued

Table 1.—Determination of factors by which to multiply scale readings on Leeds and Northrup potentiometer to obtain normal incidence radiation in gram calories—Continued

Date and hour		Scale	Factor	Scale	Factor	1	Means		Date and hour		Scale	Factor	Scale	Factor		Means	
angle	Q	(Ep- pley)	Q/Ep. pley	(Clark)	Q/ Clark	Q	Eppley	Clark	angle	Q 	(Ep- pley)	Q/Ep- pley	(Clark)	Q/ Clark	Q	Eppley	Clark
1939 March 7:			1						1939 March 7:								
3:18	1. 163 1. 168 1. 158 1. 160 1. 174	46. 8 47. 1 47. 0 47. 0 47. 2	249 248 246 247 249	89. 4 89. 5 89. 3 89. 6 89. 4	. 0130 131 130 130 131				1:45	1. 371 1. 363 1. 345 1. 333 1. 323	53. 9 53. 9 53. 4 53. 0 52. 9	254 253 252 252 252 250			1. 347	.02522	
3:09	1. 192 1. 192 1. 170 1. 176 1. 176	47. 8 47. 8 47. 4 47. 5 47. 5	249 249 247 247 247	89. 2 89. 2 89. 5 89. 5 89. 6	134 134 131 131 131	1, 173	. 02478	. 01313	2.00	1. 322 1. 326 1. 327 1. 329	52. 9 53. 0 53. 0 53. 0	250 250 250 251			1. 326	. 02502	
2:43	1, 289 1, 290 1, 285 1, 276 1, 267 1, 273 1, 274	51. 0 51. 1 51. 1 51. 4 51. 8 51. 7 51. 9 51. 9	248 252 252 250 246 245 245 245 245 246	96. 1 96. 2 96. 3 96. 3 96. 3 96. 6 97. 0 96. 8	131 134 134 133 133 131 131 131 132 131		. 02476		2:37	1. 261 1. 267 1. 253 1. 240 1. 235 1. 239 1. 246 1. 247 1. 244 1. 232	50. 7 50. 7 50. 4 50. 1 50. 0 50. 0 50. 1 50. 1 49. 8 49. 8	249 250 249 248 247 248 249 249 249 248 247	93. 4 93. 4 93. 0 92. 7 92. 8 92. 8 92. 7 92. 4 92. 4 92. 6	135 135 134 133 134 134 134 135 135 133			
1:48	1, 367 1, 371 1, 378 1, 368	54. 5 54. 5 54. 5 54. 5	251 252 253 252	101. 0 101. 5 102. 5 102. 0	135 135 135 134				2:50 March 8:	1. 240 1. 240 1. 233	50. 0 50. 0 40. 8	248 248 248	92. 6 92. 6 92. 5	134 134 133	1. 243	. 02482	. 01337
1:39	1. 361 1. 363 1. 364 1. 366 1. 371 1. 378	54. 0 54. 5 54. 5 54. 5 54. 5 54. 5	252 254 250 251 252 253	101. 0 101. 0 101. 0 101. 5 101. 5 102. 0	135 135 135 135 135 135	1. 369	. 02520	. 01349	1:32	1, 389 1, 384 1, 394 1, 415 1, 405 1, 395 1, 394 1, 375	54. 5 54. 5 55. 0 56. 0 55. 5 55. 0 55. 0	255 254 253 253 253 254 254 254 255	104. 0 104. 0 105. 5 105. 0 105. 0 104. 0 104. 0	134 133 135 135 134 134 134			
1;20	1. 431 1. 426 1. 410 1. 422	55. 5 56. 5 56. 0 55. 5 56. 0	255 253 255 254 254	106. 0 106. 0 106. 0 106. 0 105. 0	134 135 135 133 136				1:23 0:54	1. 389 1. 391 1. 378 1. 379	55. 0 55. 0 54. 5 54. 5	253 253 253 253	104. 0 104. 0 103. 0 103. 0	134 134 134 134	1. 393	, 02537	.01337
1:11	1. 414 1. 415 1. 421 1. 425	55. 5 55. 5 56. 0 56. 5	255 255 254 252	105. 0 105. 0 105. 0 105. 0	135 135 136 136	1. 420	. 02541	. 01348	·	1. 378 1. 381 1. 376 1. 368 1. 381 1. 390	54. 5 54. 5 54. 5 54. 0 54. 0 55. 0	253 253 253 254 256 253	103. 0 103. 0 103. 0 103. 0 103. 0 103. 0	134 134 134 133 134 135			
0:59	1. 410 1. 403 1. 404 1. 433 1. 441	55. 0 55. 0 55. 0 45. 0 56. 0	256 255 255 256 256 257	105. 0 105. 0 105. 0 105. 0 105. 0	137 138				0:44 0:28	1. 387 1. 381	55. 0 55. 0 49. 9	252 251 251	103. 0 103. 0 103. 0	135 134 . 0133	1. 380	. 02531	. 01341
0:50	1. 448 1. 433 1. 420 1. 414 1. 411	56. 0 56. 0 55. 5 55. 5 55. 6	258 256 256 256 255 257	105. 0 105. 0 105. 0 105. 0 105. 0	138 137 136 135 135		. 02581		0:39	1. 214 1. 232 1. 222 1. 208 1. 203 1. 191 1. 180 1. 200 1. 214	44. 9 50. 0 49. 0 49. 2 48. 2 48. 1 48. 0 49. 0	243 246 250 248 250 248 248 246 245 245	92. 8 93. 0 92. 8 92. 2 91. 2 90. 6 90. 2 91. 6	133 132 133 131 131 132 131 131 131	1. 208		. 01319
0:45	1. 436 1. 420 1. 423 1. 417 1. 422 1. 431	56. 0 55. 5 55. 5 55. 5 55. 5 56. 0	. 0256 256 256 255 256 256 256	106. 5 106. 0 106. 0 105. 5 106. 0 106. 5	134 134 135 134				1:14	1. 194 1. 186 1. 182 1. 174 1. 189	47. 8 47. 6 47. 6 47. 6 48. 0	250 249 248 247 248	88. 0 88. 0 88. 4 89. 2 89. 0	136 135 134 132 134	1. 185		.01342
0:36	1. 432 1. 447 1. 439 1. 427	56. 0 56. 5 56. 5 56. 0	256 256 256 255 255 255	107. 0 107. 0 107. 0 106. 5	134 135 135 134			. 01344	1:58	1. 143 1. 150 1. 133 1. 137	46. 5 46. 0 46. 0 46. 0	246 246 246 247	86. 2 86. 2 86. 4 86. 8	131 131			
0:31	1, 481 1, 441 1, 414 1, 441 1, 439	57. 5 56. 5 55. 5 56. 5 56. 5	257 255 255 255 255 265	108. 0 106. 0 104. 0 106. 0 106. 0	138 136 136 136 136				2:07	1. 137 1. 133 1. 126 1. 117 1. 114 1. 111	46. 0 45. 9 45. 8 45. 6 45. 4 45. 4	247 247 246 245 245 245 245	86. 8 86. 5 86. 2 85. 2 85. 2 84. 2	131 131	1.130		.01313
0:40	1. 420 1. 414 1. 407 1. 396 1. 388	56. 0 55. 5 55. 0 55. 0 54. 5	255 254 256 256 254 255	105. 0 105. 0 104. 0 104. 0 104. 0	135 135 135 135 134			.01356	2:30	1. 153 1. 148 1. 136 1. 133 1. 136 1. 133	46. 4 46. 2 46. 1 46. 0 46. 0 46. 0	248 248 246 246 247 246	85. 6 85. 6 85. 6 85. 6 85. 5 85. 5	133 133 133			
1:45	1. 361 1. 361 1. 366 1. 366 1. 369	54. 0 54. 0 54. 0 54. 0 54. 0	252 252 253 253 254						2:40	1. 123 1. 111 1. 108 1. 110	46. 0 45. 8 45. 3 45. 3	244 243 245 245	85. 2 85. 0 84. 8 84. 8	132 131 131 131			.01326
:	1. 360 1. 328 1. 323 1. 338 1. 359	53. 9 53. 2 53. 2 53. 3 53. 6	252 250 249 251 254	off S	cale	1.342	. 02512		3:07	1.077 1.060 1.046 1.047 1.041	43. 9 43. 6 43. 0 42. 8 49. 6	245 243 243 245 245	82. 2 82. 0 81. 7 81. 7 81. 6	131 129 128 128 128	1		.01288

Table 1.—Determination of factors by which to multiply scale readings on Leeds and Northrup potentiometer to obtain normal incidence radiation in gram calories—Continued

Table 1.—Determination of factors by which to multiply scale readings on Leeds and Northrup potentiometer to obtain normal incidence radiation in gram calories—Continued

Date and hour	Q	Scale	Factor	(Clark)	Factor Q/	Means				
angle	.	(Ep- pley)	Q/Ep- pley)		Clark	Q	Eppley	Clark		
1939										
March 9: 3:22	0. 860 0. 867	35. 8 36. 1	240 240	67. 8 68. 2	127 127					
	0.852 0.860 0.864 0.857 0.854	36. 1 35. 8 35. 8	238 240	68.0 68.4	126 126					
	0.864 0.857	36.0 35.8	240 239	68.7 68.6	126 125					
	0.854 0.860	35, 8 36, 0	238 239	68. 4 69. 4	125 124					
	0.864 0.866	36. 2 36. 2	239 239	69. 6 69. 7	124 124	0.860	. 02392	, 01250		
	0.912	37.7 37.9 38.0	242 245	71.0 71.9	129 129					
	0. 927 0. 927 0. 931	38.0	244	71.9 72.1	129 129	1.	i			
		38. 4 38. 8	243 242	72.1 74.0	130 127					
	0. 938 0. 938 0. 944 0. 952 0. 957 0. 961 0. 968	39.0° 39.3	241 240	74.0	128					
	0.952 0.957	39.8 39.8	239 240	74. 5 74. 8 76. 0	128 128) 		
	0.961 0.968	40. 1 40. 4	240 240	76.0 75.4	126° 128					
		40.6 40.8	241 241	75.4 75.6	130 130					
	0. 983 0. 982 0. 980	40.5	242	76.9	128 127					
	0 983	40. 4 40. 1	243 245	77.0 77.2 77.4	127	0.050		01005		
2:51	0. 996	40.8	244	77.4	129	0, 956	. 02419	, . 01285		
2:50	0.998 0.994	40.8	. 0245	75. 8 76. 0	. 0132 131					
	1.002	41. 0 41. 0	243 244	76. 4 77. 0	131 131					
	1.009 1.009	41.0 41.1 41.2	246 245	1 77.0 1	131					
j	1.005 1.002	41.0	244 244	77. 4 77. 4	130 129					
	1,002 1	40. 9 40. 8	245	77. 0 76. 7	130 131					
. 1	1. 004 1. 007 1. 011	40.9 41.4	246 246 244	76. 5 77. 2	132 131		l. 			
	1. 02 5 1. 03 7	41.8 42.0	245	77. 6 77. 6	132 134					
	1.033	42.0	247 246	78. 4 78. 4	132 132					
j	1. 036 1. 036	42. 2 42. 4	245 245	78.2	132					
	1.030 1.035	42. 2 42. 3	244 245	78. 8 79. 2	131 130					
	1. 039 1. 042	42. 4 42. 4	245 246	78. 8 78. 3	132 133					
j	1. 039 1. 043	42. 6 42. 8	244 244	78. 3 78. 2	133 133					
	1. 047 1. 052	42.9 42.8	244 246	79. 2 78. 7	132 134					
	1. 052	42.9	245	79.4	132	1.024	. 02449	. 01316		
	1. 055 1. 058	43. 2 43. 1	244 246	78.8 78.6	134 135					
,	1. 061 1. 062	43. 4 43. 4	244 244	78. 9 78. 9	134 135					
	1. 059 1. 058	43.3 43.3	245 244	79. 3 79. 8	134 133					
•	1.056 1.054	43. 2 43. 1	244 245	80. 4 80. 5	131 131					
	1.057	43. 1	245	80. 6 81. 1	131 130					
	1. 067 1. 076	43. 0 43. 4	248 248	81. 6 81. 8	132	,				
	1. 070 1. 067	43.5 43.5	246 245	82.4	131 130					
	1. 070 1. 070	43. 6 43. 7	245 245	82. 5 82. 6	130 130					
	1. 073 1. 077	43. 8 44. 0	245 245	82. 7 82. 8	130 130					
	1. 080 1. 081	44. 2 44. 4	244 243	82. 4 81. 7	131 132					
	1.086	44.5	244	81. 9 82. 4	133 133					
	1. 093 1. 099	44.7 44.8	245 245	82.4	133					
	1. 103 1. 103	45. 1 45. 2	245 244	82. 8 82. 9	133	}				
2:00	1. 102	45. 2	244	82. 9	133	1.073	. 02449	. 01324		
1:59	1.099 1.096	45. 2 45. 0	243 244	83. 8 83. 3	131 132					
ļ	1. 093 1. 089	45. 2 44. 8	242 243	83. 9 84. 3	130 129					
	1.090	44.5	245	84. 6 83. 8	129 131					
	1. 091 1. 101	44. 6 44. 7	245 246	84.4	130			01000		
1:52	1. 107	44.8	247	84.9	130	1.096	.02444	. 01302		
			1		100	1	ı			
1:22	1. 122 1. 127	45. 1 45. 1	249 250	84.4 84.4	133 134					
1:22	1, 122 1, 127 1, 127 1, 121	45. 1 45. 1 45. 0 45. 6	249 250 250 246	84. 4 84. 4 84. 6 84. 7	134 133 133 132					

Date and hour	Q (H	Scale	Scale (Ep- pley) Factor Q/Ep- pley)	Scale	Factor	Means				
angle				(Clark)	Q/ Clark	Q	Eppley	Clark		
1939										
March 9: 1:22	1. 144 1. 126 1. 118 1. 117 1. 122 1. 140 1. 145 1. 140 1. 145 1. 148 1. 152 1. 159 1. 162 1. 163	45. 8 46. 1 45. 8 45. 8 46. 1 46. 2 46. 3 46. 2 46. 4 46. 5	250 244 244 244 245 247 248 247 248 249 250 250	84. 8 84. 9 85. 4 85. 2 84. 9 86. 2 86. 0 84. 9 85. 7 85. 8 85. 7 85. 8	135 131 131 132 133 133 135 134 134 135 136 136		.02478			
0:52.	1. 150 1. 151 1. 147 1. 146 1. 131 1. 123 1. 135 1. 132 1. 116 1. 108	46. 2 46. 1 46. 0 45. 8 45. 3 45. 6 45. 4 45. 2 45. 1	0249 250 249 249 250 246 250 249 248 246	85. 6 84. 3 84. 1 84. 6 84. 5 83. 2 83. 0 83. 0 82. 4 82. 2	. 0134 136 135 134 135 136 136 136					
0:37	1. 090 1. 084 1. 075 1. 070 1. 071 1. 065 1. 062 1. 058 1. 051 1. 050	43.8 43.8 43.4 43.3 43.2 43.2 43.1 43.1	249 248 246 247 247 247 246 245 244 244	79.8 79.6 79.6 79.6 79.3 79.2 79.1 79.0 79.0	137 136 135 134 135 134 134 134 133 133		. 02463			
1:20	0. 977	40.0	244	74.0	132	0. 977	. 02440	. 013		
March 10: 3:27	0. 887 0. 884 0. 887 0. 890 0. 901 0. 898 0. 906 0. 910 0. 910 0. 901 0. 903 0. 903 0. 903 0. 905	36. 0 36. 7 36. 9 37. 0 37. 4 37. 4 37. 3 37. 0 37. 2 37. 3 37. 3 37. 3 37. 3 37. 3 37. 3 37. 3	241 241 242 241 243 240 239 240 243 247 245 244 241 242 241 241 243 244 244 243	67. 4 67. 6 67. 7 68. 6 68. 7 69. 0 69. 0 69. 8 70. 0 69. 8 69. 8 69. 8 69. 8 69. 8 69. 2 69. 2	132 131 131 131 130 130 130 130 130 130 129 129 130 130 130 131					
3:06	0. 929 0. 945 0. 957 0. 953 0. 944 0. 947 0. 946 0. 957 0. 978 0. 987	37. 9 38. 8 39. 0 39. 2 38. 8 38. 8 38. 8 39. 2 39. 8	245 244 246 243 243 244 244 244 244 246 248	71. 2 72. 0 73. 4 73. 4 73. 3 72. 8 73. 0 73. 4 74. 8	131 130 130 129 130 130 130 130 131 131	0.954		.013(

Table 2.—Probable errors of individual readings, and of means from lines of best fit and from means of all

		le errors: ⊢1.45 gr.•			: Range		
•	Potent	iometer	Clark with	Potentiometer		Clark with micro-	
	Eppley	Clark	micro- amme- ter	Eppley	Clark	amme- ter	
Means from line (variable factor) Individual readings from line (vari-	±0. 24	±0.84	±1.81	±0.40	±1.40	±3.02	
able factor)	±0.37	±0.92		±0.62	±1.53		
Means from mean of all (constant factor)	±1.18	±1.31	±2.69	±1.97	±2.18	±4.48	

These small probable errors from the line of best fit unquestionably show the need for using variable factors rather than a constant factor. In fact they are as close as would be expected with the regular substandard instruments for the following reasons:

(1) Normal incidence radiation never is uniform at sea level, owing chiefly to turbidity. Waviness therefore in a normal incidence curve, although slight on the best of

days, is natural

(2) In the operation of the Smithsonian silver-disk pyrheliometer the shutter is open for 2 minutes, then closed for 2 minutes. The alternations with the Marvin resistance pyrheliometer occur every minute. It is conceivable under adverse conditions that the shutter might be open during low radiation receipt, or that these conditions might be reversed. It is obvious that an error is introduced when comparing such an instrument against one that gives instantaneous and continuous readings.

(3) Owing to the personal equation it is necessary for each user of a Smithsonian or a Marvin pyrheliometer to personally read the instrument when checking against Smithsonian standards at the Astrophysical Observatory. A change of observers of necessity introduces another

small source of error.

(4) While the design of all instruments here mentioned calls for an angular opening of 5°43′, in practical construction it mechanically is impossible to adhere to these measurements perfectly. As the annulus about the sun is by far the brightest portion of the sky, any increase or dimunition of the diameter of this annulus, even very slight, creates an error which is particularly appreciable on hazy days.

(5) The addition of a highly polished thin quartz or glass window over the receiving end of the Eppley and Clark pyrheliometers changes the spectral distribution of energy received on the thermoelectric surfaces enough to

produce another small error.

(6) The receiving surfaces of all the pyrheliometers, especially those without sealed windows, undergo slight changes in their absorption coefficients owing to dust and other extraneous material falling upon their surfaces.

- (7) Any recording mechanism lacks 100 percent precision owing to several factors, among which may be cited (a), nonuniform scale divisions; (b), incorrect setting of the zero and pen; (c), change in length and width of paper because of humidity variations; (d), slight changes in the e. m. f. of the standard cell used with potentiometers; (e), irregular rotation of the record roll; (f), zero shift for a number of reasons; and (g), the gradual lowering of the e. m. f. of the operating battery between checks against the standard cell.
- (8) Rapid temperature changes of the free air, and winds of appreciable velocity, vitiate slightly the readings of all phyrheliometers of the types here mentioned.

As previously stated, only those readings made after the potentiometer was thoroughly adjusted to the highest practical efficiency were used in these comparisons. After this adjustment the instrument gave extraordinarily good results as shown by a continuous record, for more than 100 hours, of the e.m. f. generated by the vacuum thermocouple when receiving its energy from a well-seasoned lamp in series with a constant voltage regulator.

In order to minimize errors of paper shrinkage and expansion, a special type of record paper is used which

has a low coefficient of expansion.

Although the potentiometer automatically balances the dry cells against the standard cell every 43 minutes, we also make this balance manually immediately preceding each series.

All the other instruments were thoroughly checked and placed in first-class condition before the calibrations. The Smithsonian silver-disk pyrheliometer was checked against Smithsonian standards at the Astrophysical Observatory, and used only twice before the tests; all instruments were realigned, and indicator points re-etched to insure their correct setting on the sun; the Marvin pyrheliometer was checked against the silver-disk; the signal-clock was regulated to run at a uniform rate; the microammeter was tested at the National Bureau of Standards and returned to the factory for replacement of a faulty bearing, after which it was calibrated at the Bureau, and a table giving the true values in microamperes of the scale readings was used to reduce the observations.

Upon first thought it might appear that the logical method of making these tests would be to run the two thermoelectric pyrheliometers against a standard artifical source of radiation. Practical limitations to date have prevented much work along this line, although some tests were made with the vacuum thermocouple at the Bureau of Standards; these were meager owing to lack of a light source of sufficient energy. Moreover, it is impossible practically to obtain a point-source of light; and as yet no artifical source of energy approximates

closely the spectral distribution of solar energy.

Attempts to insure a high degree of accuracy have in the past so complicated the apparatus and rendered it so expensive that we have had to limit sharply the number of solar observational stations. It is thought that the newer type of apparatus will relieve this situation. Without doubt the utmost in precision is required in many special researches; but in the case of the Weather Bureau, lack of personnel and equipment prohibit the general use of precision apparatus in the field, although we maintain such instruments at our central observatory, and for general radiation climatology, high precision is not necessary.

Thermoelectric pyrheliometers are especially well adapted for measuring the red and yellow components between 0.61 and 0.51 μ 3 and have been used for this purpose by both this Bureau and the Blue Hill Meteor-

ological Observatory.

Upon completion of the tests, the manufacturers immediately took steps to redesign the thermopile, particularly as to spacing of the elements, so as to decrease the variability of the factor values. While preliminary tests on one of these new pyrheliometers show a marked improvement in performance, the data obtained so far are too meager to give definite results.

³ Kimball, Herbert H. Determinations of atmospheric turbidity and watervapor content. Monthly Weather Review 64: 1-5, 1936.

Kimball, Herbert H. and Hand, Irving F. The use of glass color screens in the study of atmospheric depletion of solar radiation. Monthly Weather Review 61: 80-83, 1933.

Table 3.—Comparison between the constant factor and the variable factors of the Eppley pyrheliometer

,		NO DPP	og Fø.		
		(3)	(4)	(5)	(6)
(1) Scale ¹	(2) Factor	Gram-	calories	Percent depar- ture of	Corre- spond- ing
Scale -	ractor	(1)+(2)	0.0249 ×(1)	(4) from (3)	milli- velts
21. 0 23. 1 25. 3 27. 5 29. 6 31. 7 35. 8 37. 8 37. 8 41. 4 44. 9 46. 9	0. 0234 235 236 237 238 239 240 241 242 243 244 245 246 247 248	0. 491 . 545 . 597 . 652 . 704 . 758 . 809 . 863 . 915 . 967 1. 010 1. 056 1. 105 1. 158 1. 205	0. 523 . 575 . 630 . 685 . 737 . 789 . 839 . 891 . 991 1. 031 1. 073 1. 118 1. 168 1. 210	+6.5 +5.9 +5.5 +5.1 +4.1 +3.7 +2.8 +2.5 +2.1 +1.6 +1.2 +0.9	0. 61 . 67 . 74 . 81 . 88 . 95 1. 02 1. 08 1. 14 1. 19 1. 25 1. 31 1. 36 1. 41
50. 1 51. 4 52. 5	249 249 250 251	1. 247 1. 285 1. 318	1. 247 1. 280 1. 307	-0.4 -0.7	1. 49 1. 53 1. 57
53. 5 54. 4 55. 4 56. 2 56. 9	252 253 254 255 256	1. 348 1. 376 1. 407 1. 433 1. 457	1. 332 1. 355 1. 379 1. 399 1. 417	-1. 2 -1. 5 -2. 0 -2. 3 -2. 7	1. 61 1. 64 1. 67 1. 70 1. 72
57. 6 58. 2 59. 0 59. 7	257 258 259 260	1. 480 1. 502 1. 528 1. 552	1. 434 1. 449 1. 469 1. 487	-3. 2 -3. 5 -3. 8 -4. 2	1. 74 1. 76 1. 78 1. 80

 $^{^1\,\}rm The$ recording micromax potentiometer used for this test has a full-scale deflection of 3 millivolts; it therefore is necessary to shift to its alternate 15-milivolt circuit when the needle reaches 100 on the scale.

The probable errors of the values in column 3 do not exceed ± 0.3 percent.

Factors to reduce scale readings on potentiometer recording e. m. f. generated by Clark thermoelectric pyrheliometer

_						
Potentiometer	Factors	Gram- calories	Potentiometer	Factors	Gram- calories	
Scale readings: 1 40.5 46.9 54.7	0. 0126 127 128	0. 510 . 596 . 700	Scale readings: 81.8 88.7 98.5	132 133 134	1. 080 1. 180 1. 320	
62.0 70.0 76.3	129 130 131	. 800 . 910 1. 000	105.2 109.0 114.0	135 136 137	1. 419 1. 482 1. 562	

With potentiometer having full scale deflection of 3 millivolts it is necessary to shift to the 15-millivolt scale when the needle approaches the top of the scale.

Our conclusions are:

- (1) Provided factors are determined according to methods here described, thermoelectric pyrheliometers are excellent for laboratory use in making routine measurements with a precision as good as that obtained with a Marvin pyrheliometer, and only slightly under the precision attained with the Smithsonian silver-disk pyrheliometer.
- (2) The advantages of the use of this type of instrument are manifold; first, a saving of at least 75 percent in the observer's time; second, the readings are continuous; and third, the simplicity of the whole apparatus eliminates much of the trouble experienced with the accessories necessary for the Marvin pyrheliometer.
- (3) The vacuum type is ideal for field use when used with a portable potentiometer, especially when weight is an important factor, as for example, when measurements are desired on high, poorly accessible mountain tops, because the whole pyrheliometer weighs less than 1 pound.
- (4) The vacuum pyrheliometer assumes equilibrium within six seconds after opening the shutter; the copperconstantan type requires about 20 seconds to reach equilibrium.
- (5) The probable errors are slightly less with the non-vacuum type.
- (6) A portable precision eye-read potentiometer is recommended for field use rather than a microammeter, as the former eliminates practically all errors arising from changes in resistance of various units in the electrical circuit.

Additional comparisons made in subzero weather early in 1941 between the Smithsonian silver disk, the Clark vacuum type and the new Eppley pyrheliometers show (1) much less variation in the factors for the new Eppley pyrheliometers with widely-spaced elements, and (2) a slight free-air temperature effect; that is, all the thermo-electric pyrheliometers tested show greater efficiency with very low free-air temperatures.

ADJUSTMENT OF AIRPORT STATION-PRESSURE RECORDS TO FORMER CITY STATION ELEVATION

By W. W. REED

[Weather Bureau, Washington, D. C., January 1941]

In the installation of mercurial barometers at the airports, the tables for reduction of station pressure to sea level were based in most cases on a station elevation corresponding exactly, or very nearly, to the elevation of the ivory point of the barometer, or to the level 8 feet above the landing field. In only a relatively few cases was the adopted station elevation made to coincide with the station elevation at the city office.

At city offices established prior to 1900, the practice has been followed since the beginning of that year of maintaining a single "station elevation" by applying a "removal correction" whenever the barometer was moved to a different elevation from that existing on January 1, 1900, so that the "station pressures" pertained to the actual elevation as of that date. Thus the adopted "station elevation" corresponded to the actual elevation of the ivory point of the barometer at the beginning of the current century. At city offices established subsequent to January 1, 1900, the adopted "station elevation" was almost invariably the actual elevation of the barometer

when the station was first established. Under this system, records of "station pressure" at city offices have been directly comparable since the dates in question by virtue of the fact the data were pertinent to a single "station elevation."

However, where city offices were closed or consolidated with the airport stations, the changes in elevation were so considerable in many cases that it was inadvisable to attempt the employment of a "removal correction" and the airport "station elevations" were maintained.

Beginning with July 1939, and prior thereto at several stations, the records of pressure at most of the airports were made official for synoptic purposes and published in the Monthly Weather Review. This procedure introduced into the homogeneity of pressure records breaks that range in value from a few thousandths of an inch, insignificant for practical purposes, to more than 0.50 inch locally in winter. In view of the need for homogeneity in respect to elevation in the study of pressure trends, action has been taken to prepare adjustments for